Risaliti and Elvis 2010A&A...516A..89R

## AGN Structure from the Three Forms of Radiation Pressure

### Martin Elvis

Harvard-Smithsonian Center for Astrophysics

#### Atomic Features in Quasar Spectra

Should have a Unified, Simple Theory



#### Major Features of Elvis 'Funnel Wind'

#### Elvis 2000 ApJ 545, 63; 2003 astro-ph/0311436



The 3 Forms of Radiation Pressure

Structure of this talk:

- 1. Compton scattering *short*
- 2. Line driving long
- 3. Dust driving *short*
- 4. Implication short

#### **Claim:** Radiation Driving Determines Quasar Structure



# **Compton Driving**

#### Only works at or above Eddington

#### Failed Compton wind in most AGN?



### Failed Compton Winds in Sub-Edd AGNs



#### NGC1365: Risaliti et al., 2005, ApJL, 630, L129



#### **Disk Geometry Creates Shielding Gas**

Risaliti and Elvis 2010A&A...516A..89R

Spherical: no net effect



 $\tau > 1 =$  "hitchiking gas" of Murray & Chiang 1997

# **Line Driving**

### Warm Absorbers and Associated Absorbers

• Line Driving: Multiplier >100 x Compton scattering in O-stars– works far below L<sub>Edd</sub>

Castor, Abbott & Klein 1975 (CAK) Murray et al, Murray & Chiang, Proga, ...

- Widely accepted for WA and UV NALs
- 2-3 phase medium fits all cases Krongold et al., Netzer et al., Andrade-Velazquez et al. 2010ApJ...711..888A
- Equatorial?
- Wide range of radii?
- Non-analytical, hard to span parameter space with hydro simulations

#### A Non-Hydrodynamical Model for Acceleration of Line-Driven Winds in AGNs Risaliti and Elvis 2010A&A...516A..89R

- Supersonic wind from early on
- Separate acceleration from launching
  - Initial vertical velocity
- Treat as gas elements
  - BH gravity
  - Line driven radiation pressure
  - Point X-ray source
  - Disk UV source
  - Radiative transfer
  - CLOUDY + CAK = QWIND

## Results



#### Reproduces Major Features of Elvis 'Funnel Wind'

Elvis 2000 ApJ 545, 63; 2003 astro-ph/0311436





## Exploring Parameter Space BELR Wind

#### Rapid X-ray Eclipses – days to hours



#### Transiting, Transient BLR clouds



Sizes D ~  $10^{13}$  cm. Blobs or sheets?

# **Dust Driving**





Other "type changing" AGNs

## **BLR Clouds Launched by Dust Driving**

- Most efficient absorber
- Dust formed when outer disk atmosphere <1000K
- $\bullet$  Happens from H $\beta$  radius outward
- Slow N<sub>H</sub> eclipses ~years



#### Czerny & Hrniewicz 2011 A&A 525 L8

## The 3 Forms of Radiation Driving Determine Quasar Structure



## Implication: Matter Launched at all Disk Radii

#### Evidence: Short-lived BLR clouds, constantly renewed



### Evidence: "Cometary" Eclipsers



## BLR 'Breathing'

Peterson et al. 1999 ApJ, 510, 659



→ gas ejected over whole range of radii

Question becomes: What is the Launching Mechanism?

- Distinct from Acceleration
- One mechanism, or multiple?
- Multiple:
  - Dust Driving: T<sub>disk</sub> < 1000 K
  - Line Driving: T<sub>disk</sub> >~ 30,000 K
  - What about  $1000 < T_{disk} < 30,000$ ?
- Single: Magnetic recombination
  - Solar Coronal Mass Ejections reach 1000 km/s
  - Driven by MRI?
  - Is disk ionized enough at all radii?
- MHD winds?

## The 3 Forms of Radiation Driving Determine Quasar Structure



Now explore broad parameter space, launching

High L/L<sub>Edd</sub> does not automatically imply a faster wind



## The 3 Forms of Radiation Driving Determine Quasar Structure



Now explore broad parameter space, launching